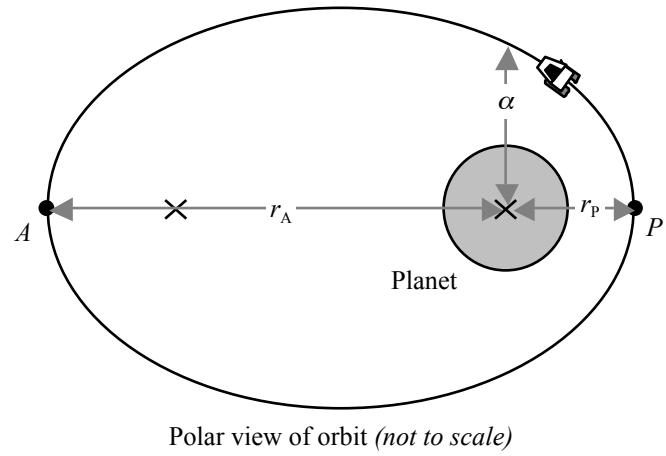


1. Consider again the orbit of the shuttlecraft shown in section II of the tutorial, reproduced below. In this problem we develop some quantitative relationships among the semi-major axis  $a$ , eccentricity  $\varepsilon$ , latus rectum  $\alpha$ , apogee distance  $r_A$ , and perigee distance  $r_P$ .

- Each focus of an ellipse is located a distance  $\varepsilon a$  away from the geometric “center” of the ellipse. Use this fact to express  $r_A$  and  $r_P$  in terms of  $a$  and  $\varepsilon$ . Show all work.
- Show that the latus rectum  $\alpha$  can be expressed in terms of  $a$  and  $\varepsilon$  as follows:  $\alpha = a(1 - \varepsilon^2)$ .

(Hint: For any point along the ellipse, how must the sum of the distances from that point to each focus relate to  $a$ ? Apply this relationship to the point at which the latus rectum is measured.)



- Combine your results from parts a and b in order to express  $r_A$  and  $r_P$  in terms of  $\alpha$  and  $\varepsilon$  (rather than in terms of  $a$  and  $\varepsilon$ , as you did in part a of this problem). Show all work.

2. In this problem we will make use of the fact that for closed orbits the latus rectum  $\alpha$  of the orbit is equal to  $\alpha = l^2/GM$ , where  $G$  is the universal gravitational constant,  $l$  is the angular momentum *per unit mass* of the orbiting body, and  $M$  is the (much larger) mass of the body at one focus of the orbit.

The following questions refer to the shuttlecraft orbit shown in Problem 1 above.

- Use your knowledge of Kepler’s laws and your results in Problem 1 to determine expressions for (i) the shuttle’s speed  $v_P$  at perigee, and (ii) the shuttle’s speed  $v_A$  at apogee. Write each expression in terms of  $G$ ,  $M$ ,  $\varepsilon$ , and the corresponding distance ( $r_P$  or  $r_A$ ).
- Use your results from part a of this problem to calculate the quantities listed below. In your work, assume that the planet has approximately the same radius and mass as Earth (i.e.,  $R_{\text{planet}} \approx R_E = 6.38 \times 10^6$  m, and  $GM \approx GM_E = 3.98 \times 10^{14}$  N·m<sup>2</sup>/kg). In addition, suppose that the orbit of the shuttle has perigee  $r_P = 2R_E$  and apogee  $r_A = 8R_E$ .
  - the eccentricity  $\varepsilon$  of the original (elliptical) orbit
  - the latus rectum  $\alpha$  of the orbit
  - the speed  $v_P$  at perigee in the orbit
  - the speed  $v_A$  at apogee in the orbit
  - the minimum speed your shuttlecraft must attain at point  $P$  (e.g., by quickly firing your thrusters as you pass that point) in order to leave the planet

(Hint: What minimum value of eccentricity  $\varepsilon$  is required for the shuttle’s orbit in this case?)

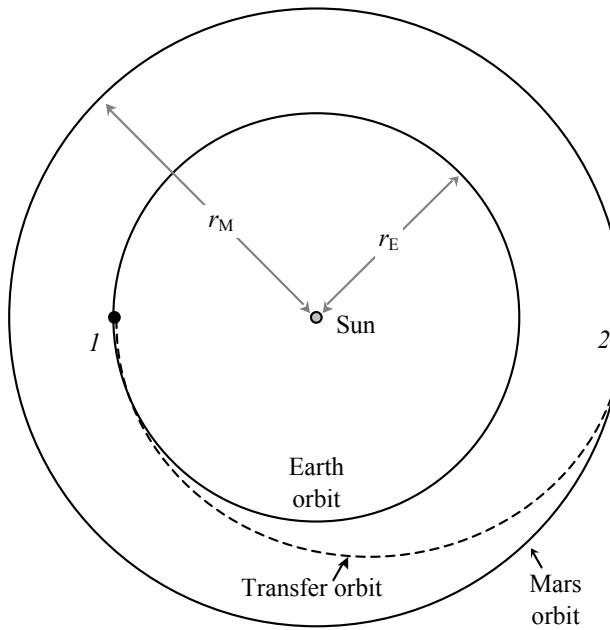
## Homework: Energy and angular momentum for closed orbits

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3. In interplanetary exploration, elliptical transfer orbits, like the one shown from Earth to Mars, allow space probes to reach the intended destination using a minimum of fuel. (Such orbits are called *Hohmann transfer orbits*.)

*Note:* Assume the orbits of Earth and Mars about the Sun are circular with radii 1.000 AU and 1.524 AU, respectively. Ignore any effects due to the rotation of the Earth.

- a. Describe *qualitatively* the maneuvers required for the probe upon entering and leaving the transfer orbit. That is, must the probe *increase or decrease* its speed (i) upon entering the transfer orbit at point 1? (ii) upon entering Mars orbit at point 2? Explain.



- b. Determine the following quantities:

- the semimajor axis, eccentricity, and latus rectum of the transfer orbit
- the time required for the probe to travel from point 1 to point 2 along the transfer orbit

(*Note:* When measuring time in units of earth-years and distance in AU's, the quantity  $4\pi/GM_{\text{sun}}$  has a value of unity.)

- c. It is said that interplanetary expeditions can be launched at only certain times, or “launch opportunities.” In light of your results from part b above, explain what is meant by this statement.
- d. Make your results from part (a) quantitative (rather than just qualitative) by calculating:
- the change in speed (in m/s) required for the probe to enter the transfer orbit at point 1
  - the change in speed (in m/s) required for the probe to enter Mars orbit upon reaching point 2

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**Homework: Energy and angular momentum for closed orbits**

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4. Consider again the pairs of orbits you compared in parts A and B in section III of the tutorial.
  - a. In part A of section III you considered two orbits in which a shuttlecraft would have the same total energy. Determine the angle formed by the orbits at a point where they intersect. Clearly show all work.

(Hint: Consider the ratio of angular momenta  $|\vec{L}_{\text{elliptical}}|/|\vec{L}_{\text{circular}}|$  evaluated at the intersection point.)
  - b. For any closed orbit of semi-major axis  $a$ , show that the speed  $v(r)$  of the shuttlecraft at any point along its orbit around the planet (mass  $M$ ) is expressed as the following function of  $r$ :

$$v(r) = \sqrt{GM\left(\frac{2}{r} - \frac{1}{a}\right)},$$

- c. In part B in section III you considered two orbits in which a shuttlecraft would have the same angular momentum. Determine the angle formed by the orbits at a point where they intersect.

(Hint: Start this part the same way you did part a of this problem, although here the ratio of angular momenta for the orbits is now equal to 1. How can you use your result in part b of this problem to help you here?)